

**Supplementary Material to  
Burnett-order constitutive relations, second  
moment anisotropy and co-existing states in  
sheared dense gas-solid suspensions**

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## Appendix A. Evaluation of collision integrals $\Theta_{\alpha\beta}$ , $A_{\alpha\beta}$ , $\hat{E}_{\alpha\beta}$ and $\hat{G}_{\alpha\beta}$

The collision integrals  $\Theta_{\alpha\beta}$  (2.14),  $A_{\alpha\beta}$  (2.16a),  $\hat{E}_{\alpha\beta}$  (2.16b) and  $\hat{G}_{\alpha\beta}$  (2.16c) appearing in the second-moment balance (2.24a-d) of the main text of Saha & Alam (2020) can be simplified to:

$$\left. \begin{aligned} 2\Theta_{xy} &= \frac{3(1+e)\rho\nu g_0 T}{\pi^{\frac{3}{2}}} \left( \cos 2\phi \mathcal{J}_{012}^{30} - \sin 2\phi \mathcal{J}_{102}^{30} \right) \\ \Theta_{xx} + \Theta_{yy} &= \frac{3(1+e)\rho\nu g_0 T}{\pi^{\frac{3}{2}}} \mathcal{J}_{002}^{30} \\ \Theta_{yy} - \Theta_{xx} &= \frac{3(1+e)\rho\nu g_0 T}{\pi^{\frac{3}{2}}} \left( \mathcal{J}_{102}^{30} \cos 2\phi + \mathcal{I}_{012}^{30} \sin 2\phi \right) \end{aligned} \right\}, \quad (\text{A } 1)$$

$$\left. \begin{aligned} A_{xx} &= -\frac{6(1-e^2)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} \left[ \frac{1}{2} (\mathcal{H}_{003}^{30} - \cos 2\phi \mathcal{H}_{103}^{30} - \sin 2\phi \mathcal{H}_{013}^{30}) \right] \\ A_{yy} &= -\frac{6(1-e^2)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} \left[ \frac{1}{2} (\mathcal{H}_{003}^{30} + \cos 2\phi \mathcal{H}_{103}^{30} + \sin 2\phi \mathcal{H}_{013}^{30}) \right] \\ A_{zz} &= -\frac{6(1-e^2)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} \mathcal{H}_{003}^{12} \\ A_{xy} &= -\frac{6(1-e^2)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} \left[ \frac{1}{2} (\cos 2\phi \mathcal{H}_{013}^{30} - \sin 2\phi \mathcal{H}_{103}^{30}) \right] \end{aligned} \right\}, \quad (\text{A } 2)$$

$$\left. \begin{aligned} \hat{E}_{xx} &= -(\hat{E}_{yy} + \hat{E}_{zz}), \\ \hat{E}_{yy} &= -\frac{6(1+e)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} \\ &\times \left[ \eta \left\{ \mathcal{H}_{101}^{31} - \mathcal{H}_{011}^{32} + \cos 2\phi \left( \mathcal{H}_{201}^{31} - \mathcal{H}_{021}^{31} + \mathcal{H}_{111}^{50} \right) - \sin 2\phi \left( \mathcal{H}_{201}^{30} + \mathcal{H}_{021}^{32} \right) \right\} \right. \\ &\quad \left. + 3\lambda^2 \left\{ \mathcal{H}_{001}^{32} + \cos 2\phi \left( \mathcal{H}_{101}^{32} + \mathcal{H}_{011}^{31} \right) + \sin 2\phi \left( \mathcal{H}_{011}^{32} - \mathcal{H}_{101}^{31} \right) \right\} \right] \\ \hat{E}_{zz} &= -\frac{12(1+e)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} \left\{ -\eta \left( \mathcal{H}_{101}^{31} - \mathcal{H}_{011}^{32} \right) - \lambda^2 \left( \mathcal{H}_{001}^{32} \right) \right\} \\ \hat{E}_{xy} &= -\frac{12(1+e)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} \left[ \frac{1}{2} \eta \left\{ \left( 2\mathcal{H}_{111}^{31} - \mathcal{H}_{021}^{32} - \mathcal{H}_{201}^{30} \right) \cos 2\phi \right. \right. \\ &\quad \left. \left. + \left( \mathcal{H}_{021}^{31} + \mathcal{H}_{111}^{32} - \mathcal{H}_{111}^{30} - \mathcal{H}_{201}^{31} \right) \sin 2\phi \right\} \right. \\ &\quad \left. + \frac{3}{2} \lambda^2 \left\{ \left( \mathcal{H}_{011}^{32} - \mathcal{H}_{101}^{31} \right) \cos 2\phi - \left( \mathcal{H}_{101}^{32} + \mathcal{H}_{011}^{31} \right) \sin 2\phi \right\} \right] \end{aligned} \right\} \quad (\text{A } 3)$$

$$\left. \begin{aligned} \hat{G}_{xx} &= -(\hat{G}_{yy} + \hat{G}_{zz}) \\ \hat{G}_{yy} &= -\frac{6(1+e)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} 2R \left\{ \mathcal{K}_{00}^{31} + \cos 2\phi (\mathcal{K}_{10}^{31} + \mathcal{K}_{01}^{30}) + \sin 2\phi (\mathcal{K}_{01}^{31} - \mathcal{K}_{10}^{30}) \right\} \\ \hat{G}_{zz} &= -\frac{6(1+e)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} \left\{ -4R\mathcal{K}_{00}^{31} \right\} \\ \hat{G}_{xy} &= -\frac{6(1+e)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} (-4R) \left\{ \frac{1}{2} \cos 2\phi (\mathcal{K}_{01}^{31} - \mathcal{K}_{10}^{30}) + \frac{1}{2} \sin 2\phi (\mathcal{K}_{10}^{31} + \mathcal{K}_{01}^{30}) \right\} \end{aligned} \right\} \quad (\text{A } 4)$$

In the above expressions,  $\mathcal{H}_{\alpha\beta\gamma}^{\delta p}$ ,  $\mathcal{J}_{\alpha\beta\gamma}^{\delta p}$  and  $\mathcal{K}_{\alpha\beta}^{\delta p}$  represent tensorial integrals defined as

$$\begin{aligned} \mathcal{H}_{\alpha\beta\gamma}^{\delta p}(\eta, R, \phi, \lambda) &\equiv \int_{\theta=0}^{2\pi} \int_{\varphi=0}^{\pi} \sin^{\alpha} 2\theta \cos^{\beta} 2\theta \sin^{\delta} \varphi \cos^p \varphi \\ &\times (1 - \eta \sin^2 \varphi \cos 2\theta + \lambda^2 (3 \sin^2 \varphi - 2))^{\frac{\gamma}{2}} \mathfrak{F}(\chi[\eta, R, \phi, \lambda; \theta, \varphi]) d\varphi d\theta, \end{aligned} \quad (\text{A } 5)$$

$$\begin{aligned} \mathcal{J}_{\alpha\beta\gamma}^{\delta p}(\eta, R, \phi, \lambda) &\equiv \int_{\theta=0}^{2\pi} \int_{\varphi=0}^{\pi} \sin^{\alpha} 2\theta \cos^{\beta} 2\theta \sin^{\delta} \varphi \cos^p \varphi \\ &\times \{1 - \eta \sin^2 \varphi \cos 2\theta + \lambda^2 (3 \sin^2 \varphi - 2)\}^{\frac{\gamma}{2}} \mathfrak{G}(\chi[\eta, R, \phi, \lambda; \theta, \varphi]) d\varphi d\theta, \end{aligned} \quad (\text{A } 6)$$

$$\begin{aligned} \mathcal{K}_{\alpha\beta}^{\delta p}(\eta, R, \phi, \lambda) &\equiv \int_{\theta=0}^{2\pi} \int_{\varphi=0}^{\pi} \sin^{\alpha} 2\theta \cos^{\beta} 2\theta \sin^{\delta} \varphi \cos^p \varphi [(1 - 2\lambda^2) \{ \sin(2\phi + 2\theta) - \cos \varphi \\ &\times \cos(2\phi + 2\theta) \} + \sin^2 \varphi \{ 3\lambda^2 \sin(2\phi + 2\theta) - \eta \sin 2\phi \}] \mathfrak{G}(\chi[\eta, R, \phi, \lambda; \theta, \varphi]) d\varphi d\theta, \end{aligned} \quad (\text{A } 7)$$

where  $\mathfrak{F}(\chi)$  and  $\mathfrak{G}(\chi)$  are given by (2.18a) and (2.18b), respectively, and each integral is evaluated over  $\theta \in (0, 2\pi)$  and  $\varphi \in (0, \pi)$ ; see figure 1 for the definitions of  $\theta$  and  $\varphi$  with reference to the contact vector  $\mathbf{k}$ . We follow an analytical approach to evaluate the above elliptic integrals following our previous works (Saha & Alam 2014, 2016) on sheared dry granular flows – the underlying methodology is identical to above works and is briefly discussed below.

Since both  $\mathfrak{F}(\chi)$  and  $\mathfrak{G}(\chi)$  are analytical functions of  $\chi$ , we substitute (2.19) into (2.18a) and (2.18b) and use the power-series representations for the complementary error function and the exponential. This yields the following infinite series representation (Saha & Alam 2014) for  $\mathfrak{F}(\chi)$  and  $\mathfrak{G}(\chi)$ :

$$\begin{aligned} \mathfrak{F}(\eta, R, \phi, \lambda; \theta, \varphi) &= -\sqrt{\pi} \left[ \frac{3R \sin^2 \varphi \cos(2\phi + 2\theta)}{\{1 - \eta \sin^2 \varphi \cos 2\theta + \lambda^2 (3 \sin^2 \varphi - 2)\}^{\frac{1}{2}}} \right. \\ &\quad \left. + \left\{ \frac{2R \sin^2 \varphi \cos(2\phi + 2\theta)}{\{1 - \eta \sin^2 \varphi \cos 2\theta + \lambda^2 (3 \sin^2 \varphi - 2)\}^{\frac{1}{2}}} \right\}^3 \right] \\ &\quad + \sum_{n=0}^{\infty} \frac{(-1)^n}{n!} \frac{3}{(2n-1)(2n-3)} \left[ \frac{2R \sin^2 \varphi \cos(2\phi + 2\theta)}{\{1 - \eta \sin^2 \varphi \cos 2\theta + \lambda^2 (3 \sin^2 \varphi - 2)\}^{\frac{1}{2}}} \right]^{2n}, \end{aligned} \quad (\text{A } 8)$$

$$\begin{aligned} \mathfrak{G}(\eta, R, \phi, \lambda; \theta, \varphi) &= \sqrt{\pi} \left[ \frac{1}{2} + \frac{4R^2 \sin^4 \varphi \cos^2(2\phi + 2\theta)}{1 - \eta \sin^2 \varphi \cos 2\theta + \lambda^2 (3 \sin^2 \varphi - 2)} \right] \\ &\quad + \sum_{n=0}^{\infty} \frac{(-1)^n}{n!} \frac{2}{(2n-1)(2n+1)} \left[ \frac{2R \sin^2 \varphi \cos(2\phi + 2\theta)}{\{1 - \eta \sin^2 \varphi \cos 2\theta + \lambda^2 (3 \sin^2 \varphi - 2)\}^{\frac{1}{2}}} \right]^{2n+1} \end{aligned} \quad (\text{A } 9)$$

Substituting (A 8-A 9) into (A 5-A 7) and carrying out term-by-term integrations over

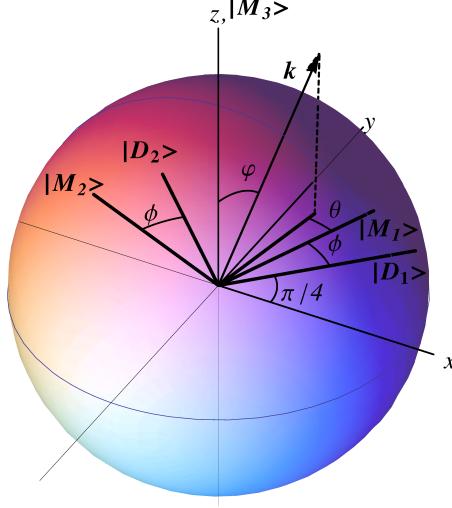


FIGURE 1. Sketch of the coordinate system  $(x, y, z)$ , with the shear-field  $[u_x(y) = \dot{\gamma}y]$  being applied on the  $(x, y)$ -plane and  $z$  is the mean vorticity direction; adapted from Saha & Alam (2016). The eigen-directions of the strain-rate tensor  $\mathbf{D}$  are marked as  $(|D_1\rangle, |D_2\rangle, |D_3\rangle)$  and that of the second moment tensor  $\mathbf{M}$  as  $(|M_1\rangle, |M_2\rangle, |M_3\rangle)$ . Note that  $\mathbf{k}$  represents the contact vector between two colliding particles, that makes an angle  $\varphi$  with  $z$ -direction and its projection on the  $(x, y)$  shear-plane makes an angle  $\theta$  with  $|M_1\rangle$  of  $\mathbf{M}$ . The principle axes frame  $(x', y', z')$  is given by the orthonormal triad of  $(|M_1\rangle, |M_2\rangle, |M_3\rangle)$  which represents a rotation of the original  $(x, y, z)$  axes by an angle  $\vartheta = \phi + \pi/4$  about the  $z$ -axis.

$\theta \in (0, 2\pi)$  and  $\varphi \in (0, \pi)$ , we obtain infinite series expressions for  $\mathcal{H}_{\alpha\beta\gamma}^{\delta p}$ ,  $\mathcal{J}_{\alpha\beta\gamma}^{\delta p}$  and  $\mathcal{K}_{\alpha\beta\gamma}^{\delta p}$  as functions of  $\eta$ ,  $\sin \phi$ ,  $\lambda$  and  $R$ .

Truncating the infinite series representations of  $\mathcal{H}_{\alpha\beta\gamma}^{\delta p}$ ,  $\mathcal{J}_{\alpha\beta\gamma}^{\delta p}$  and  $\mathcal{K}_{\alpha\beta\gamma}^{\delta p}$  at the Burnett order [i.e. the second-order in  $O(\eta^i \lambda^j R^k \sin^l 2\phi$ , with  $i + j + k + l \leq 2$ ], the following Burnett-order expressions are obtained:

$$\frac{1}{2} \left( \cos 2\phi \mathcal{J}_{012}^{30} - \sin 2\phi \mathcal{J}_{102}^{30} \right) = -\frac{4\pi}{15} \left( 8R + \sqrt{\pi}\eta \cos 2\phi \right), \quad (\text{A } 10)$$

$$\mathcal{J}_{002}^{30} = \frac{4\pi}{105} \left\{ \sqrt{\pi} \left( 35 + 96R^2 + 14\lambda^2 \right) + 48R\eta \cos 2\phi \right\}, \quad (\text{A } 11)$$

$$(\mathcal{J}_{102}^{30} \cos 2\phi + \mathcal{I}_{012}^{30} \sin 2\phi) = -\frac{8}{15} \pi^{\frac{3}{2}} \eta \sin 2\phi, \quad (\text{A } 12)$$

$$\begin{aligned} & \left[ \frac{1}{2} \left( \mathcal{H}_{003}^{30} - \cos 2\phi \mathcal{H}_{103}^{30} - \sin 2\phi \mathcal{H}_{013}^{30} \right) \right] \\ &= \frac{2\pi}{105} \left( 70 + 9\eta^2 + 42\lambda^2 + 288R^2 + 72\sqrt{\pi}\eta R \cos 2\phi + 42\eta \sin 2\phi \right), \end{aligned} \quad (\text{A } 13)$$

$$\begin{aligned} & \left[ \frac{1}{2} \left( \mathcal{H}_{003}^{30} + \cos 2\phi \mathcal{H}_{103}^{30} + \sin 2\phi \mathcal{H}_{013}^{30} \right) \right] \\ &= \frac{2\pi}{105} \left( 70 + 9\eta^2 + 42\lambda^2 + 288R^2 + 72\sqrt{\pi}\eta R \cos 2\phi - 42\eta \sin 2\phi \right), \end{aligned} \quad (\text{A } 14)$$

$$\mathcal{H}_{003}^{12} = \frac{2\pi}{105} \left( 70 + 3\eta^2 - 84\lambda^2 + 96R^2 + 24\sqrt{\pi}\eta R \cos 2\phi \right), \quad (\text{A } 15)$$

$$\left[ \frac{1}{2} \left( \cos 2\phi \mathcal{H}_{013}^{30} - \sin 2\phi \mathcal{H}_{103}^{30} \right) \right] = -\frac{4\pi}{5} \left( 2\sqrt{\pi}R + \eta \cos 2\phi \right), \quad (\text{A } 16)$$

$$\hat{E}_{yy} = -\frac{6(1+e)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} \left[ \frac{8\pi}{105} \left( \eta^2 + 21\lambda^2 + 6\sqrt{\pi}\eta R \cos 2\phi - 21\eta \sin 2\phi \right) \right], \quad (\text{A } 17)$$

$$\hat{E}_{zz} = -\frac{12(1+e)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} \left[ -\frac{8\pi}{105} \left( \eta^2 + 21\lambda^2 + 6\sqrt{\pi}\eta R \cos 2\phi \right) \right], \quad (\text{A } 18)$$

$$\hat{E}_{xy} = -\frac{12(1+e)\rho\nu g_0 T^{\frac{3}{2}}}{\sigma\pi^{\frac{3}{2}}} \left[ -\frac{4\pi}{5} \eta \cos 2\phi \right], \quad (\text{A } 19)$$

$$2R \left\{ \mathcal{K}_{00}^{31} + \cos 2\phi \left( \mathcal{K}_{10}^{31} + \mathcal{K}_{01}^{30} \right) + \sin 2\phi \left( \mathcal{K}_{01}^{31} - \mathcal{K}_{10}^{30} \right) \right\} = \frac{128\pi R^2}{105}, \quad \mathcal{K}_{00}^{31} = \frac{64\pi R}{105}, \quad (\text{A } 20)$$

$$\frac{1}{2} \cos 2\phi \left( \mathcal{K}_{01}^{31} - \mathcal{K}_{10}^{30} \right) + \frac{1}{2} \sin 2\phi \left( \mathcal{K}_{10}^{31} + \mathcal{K}_{01}^{30} \right) = \frac{2\pi}{105} \left[ \sqrt{\pi} \left( 21 + 32R^2 \right) - 16\eta R \cos 2\phi \right]. \quad (\text{A } 21)$$

### A.1. Stress tensor up-to fourth-order

The diagonal components of the dimensionless stress tensor,  $P_{ii}^* = \tilde{P}_{ii}/\rho_p \nu U_R^2$ , with  $U_R = \dot{\gamma}\sigma/2$ , correct up-to  $\mathcal{O}(\eta^i \lambda^j R^k \sin^l 2\phi, i+j+k+l \leq 4)$ , have following expressions:

$$\begin{aligned} \frac{P_{xx}^*}{T^*} &= (1 + \lambda^2 + \eta \sin 2\phi) + \frac{2(1+e)\nu g_0}{35} \left[ (35 + 96R^2 + 14\eta \sin 2\phi + 14\lambda^2) \right. \\ &\quad \left. + \frac{8}{\sqrt{\pi}} \eta R \cos 2\phi \left\{ 3(66 + 5\eta^2 - 22\lambda^2) - 160R^2 - 22\eta \sin 2\phi \right\} \right], \end{aligned} \quad (\text{A } 22)$$

$$\begin{aligned} \frac{P_{yy}^*}{T^*} &= (1 + \lambda^2 - \eta \sin 2\phi) + \frac{2(1+e)\nu g_0}{35} \left[ (35 + 96R^2 - 14\eta \sin 2\phi + 14\lambda^2) \right. \\ &\quad \left. + \frac{8}{\sqrt{\pi}} \eta R \cos 2\phi \left\{ 3(66 + 5\eta^2 - 22\lambda^2) - 160R^2 + 22\eta \sin 2\phi \right\} \right], \end{aligned} \quad (\text{A } 23)$$

$$\begin{aligned} \frac{P_{zz}^*}{T^*} &= (1 - 2\lambda^2) + \frac{2(1+e)\nu g_0}{35} \left[ (35 + 32R^2 - 28\lambda^2) \right. \\ &\quad \left. + \frac{8}{\sqrt{\pi}} \eta R \cos 2\phi (66 + 3\eta^2 - 32R^2) \right]. \end{aligned} \quad (\text{A } 24)$$

where  $T^* = \tilde{T}/U_R^2$  is the dimensionless temperature.

To calculate the fourth-order viscosity, we need to retain terms up-to the fifth-order  $\mathcal{O}(\eta^i \lambda^j R^k \sin^l 2\phi, i+j+k+l \leq 5)$  in the shear stress. The dimensionless shear stress,  $P_{xy}^* = \tilde{P}_{xy}/\rho_p \nu U_R^2$ , with  $U_R = \dot{\gamma}\sigma/2$ , at  $\mathcal{O}(R^5)$ , can be written as :

$$\begin{aligned} \frac{P_{xy}^*}{T^*} &= -\eta \cos 2\phi - \frac{4(1+e)\nu g_0}{105\sqrt{\pi}} \left[ 21R \left\{ 8 + \sqrt{\pi} \frac{\eta \cos 2\phi}{R} \right\} \right. \\ &\quad \left. + 4R^3 \left\{ 32 + 12 \frac{\lambda^2}{R^2} - \frac{\eta^2}{R^2} (2 + \cos 4\phi) \right\} \right. \\ &\quad \left. + \frac{R^5}{143} \left\{ -5120 + 15 \frac{\eta^2}{R^2} \left( 64 - 5 \frac{\eta^2}{R^2} \right) (3 + 2 \cos 4\phi) \right. \right. \\ &\quad \left. \left. + 52 \frac{\lambda^2}{R^2} \left( -128 - 33 \frac{\lambda^2}{R^2} + 12(2 + \cos 4\phi) \frac{\eta^2}{R^2} \right) \right\} \right]. \end{aligned} \quad (\text{A } 25)$$

## Appendix B. Expressions for coefficients $a_i$ in Eqn. (3.1)

The expressions for the coefficients of the temperature equation (3.1) are given by:

$$a_0 = 10000(1+e)^5\pi^{3/2}St_d^3[63(1-e)\pi - 4(11+e)St_d]^2(4+St_d^2)\nu^3g_0^3 \quad (\text{B } 1)$$

$$\begin{aligned} a_1 = & -2000(1+e)^2\pi St_d\nu g_0 \left( 18522000(1-e)\pi^2 + 463050(1-e)\pi^3 St_d \{5 - 2(1+e)(1-3e)\nu g \right. \\ & + 29400\pi^2 St_d^2 \{5(67-55e) + 2(1+e)(23+4e-3e^2)\nu g \} \\ & + 441\pi^2 St_d^3 [125(13-9e)\pi - 50(1+e)(11-54e+27e^2)\pi\nu g \\ & + 40(1+e)^2(1+3e)\{6(1-e) - \pi(1-3e)\}\nu^2g_0^2 - 864(1-e)^2(3-e)(1+e)^4\nu^3g_0^3] \\ & + 56\pi St_d^4 [1750(13-9e)\pi + 25(1+e)(277+54e-27e^2)\pi\nu g \\ & - 20(1+e)^2(1+3e)(6(11+e) - (17+3e)\pi)\nu^2g_0^2 + 648(1-e)(3-e)(1+e)^4(11+e)\nu^3g_0^3] \\ & + 12(1+e)^2 St_d^5 \nu^2 g_0^2 [2205(1-e)(1+3e)\pi^2 \\ & - (3-e)(1+e)^2 \{64(11+e)^2 + 3969(1-e)^2\pi^2\}\nu g] \\ & \left. - 336(1+e)^2(11+e)\pi St_d^6 \nu^2 g_0^2 \{5(1+3e) - 9(1-e)(3-e)(1+e)^2\nu g\} \right) \end{aligned} \quad (\text{B } 2)$$

$$\begin{aligned} a_2 = & -600(1+e)\sqrt{\pi} St_d\nu g_0 \left( 54022500(1-e)(1+3e)\pi^3 \right. \\ & + 444528000(1-e)(3-e)(1+e)^2\pi^2 St_d\nu g + 14700\pi^3 St_d^2 \{175(11-9e)(1+3e) \\ & + 5(1+e)(1829-381e-1773e^2+549e^3)\nu g + 126(1-3e)(1-e)(1+e)^3(-29+9e)\nu^2g_0^2\} \\ & + 8400(1+e)^2\pi^2 St_d^3 \nu g_0 \{5(11399-13078e+3071e^2) \\ & + 14(1+e)(541-45e-133e^2+21e^3)\nu g\} + 147\pi^2 St_d^4 [1875(13-9e)(1+3e)\pi \\ & + 125(1+e)(1351+237e-999e^2+243e^3)\pi\nu g - 50(1+e)^2\{16(1+3e)^2 \\ & + (1117-4836e-1566e^2+3636e^3-783e^4)\pi\}\nu^2g_0^2 \\ & + 1440(3-e)(1+e)^4(1+3e)\{6(1-e) - \pi(1-3e)\}\nu^3g_0^3 \\ & - 15552(1-e)^2(3-e)^2(1+e)^6\nu^4g_0^4] + 28(1+e)^2\pi St_d^5 \nu g_0 [125(705-221e)(13-9e)\pi \\ & + 50(1+e)(6417-17e-1689e^2+153e^3)\pi\nu g \\ & - 320(3-e)(1+e)^2(1+3e)\{6(11+e) - (17+3e)\pi\}\nu^2g_0^2 \\ & + 5184(1-e)(3-e)^2(1+e)^4(11+e)\nu^3g_0^3] - 24(1+e)^2 St_d^6 \nu^2 g_0^2 [1225(1+3e)^2\pi^2 \\ & - 4410(3-e)(1-e)(1+e)^2(1+3e)\pi^2\nu g_0 + (3-e)^2(1+e)^4\{64(11+e)^2 \\ & + 3969(1-e)^2\pi^2\}\nu^2g_0^2] \left. \right) \end{aligned} \quad (\text{B } 3)$$

$$\begin{aligned} a_3 = & +2520\pi \left( 96468750(1-e)\pi^3 + 91875\pi^2 St_d^2 [25(13-9e)\pi + 80(1+e)(3-8e+6e^2)\pi\nu g_0 \right. \\ & - 28(1-e)(1+e)^2\{36(3-e)(1+3e) + (1-3e)^2\pi\}\nu^2g_0^2] \\ & - 317520000(1-e)(3-e)^2(1+e)^4\pi St_d^3 \nu^3 g_0^3 \\ & + 350(1+e)\pi^2 St_d^4 \nu g_0 [125(11-24e)(13-9e)\pi \\ & - 25(1+e)\{12(1+3e)(937-1073e+252e^2) + (47-645e+1485e^2-567e^3)\pi\}\nu g_0 \\ & - 20(1+e)^2\{6(5333-2151e-4815e^2+2991e^3-462e^4) + 7(1-3e)^2(1+3e)\pi\}\nu^2g_0^2 \\ & + 12096(1-3e)(7-2e)(1-e)(3-e)(1+e)^4\nu^3g_0^3] \\ & - 800(3-e)(1+e)^4\pi St_d^5 \nu^3 g_0^3 [5(51851-59290e+13775e^2) \\ & + 126(1+e)(5+3e)(53-32e+3e^2)\nu g_0] \\ & - 42(1+e)^2\pi St_d^6 \nu^2 g_0^2 [625(13-9e)(19-6e)(1+3e)\pi \\ & + 250(1+e)(1976+129e-1557e^2+615e^3-75e^4)\pi\nu g_0 \\ & \left. - 100(3-e)(1+e)^2\{16(1+3e)^2 + (523-2316e-882e^2+1692e^3-297e^4)\pi\}\nu^2g_0^2 \right) \end{aligned}$$

$$\begin{aligned}
& +1440(3-e)^2(1+e)^4(1+3e)\{6(1-e)-(1-3e)\pi\}\nu^3g_0^3 \\
& -10368(1-e)^2(3-e)^3(1+e)^6\nu^4g_0^4] \\
& -8(3-e)(1+e)^4St_d^7\nu^3g_0^3[125(4709-4790e+977e^2)\pi \\
& +50(1+e)(3093+443e-1149e^2+45e^3)\pi\nu g_0 \\
& -160(3-e)(1+e)^2(1+3e)\{6(11+e)-(17+3e)\pi\}\nu^2g_0^2 \\
& +1728(1-e)(3-e)^2(1+e)^4(11+e)\nu^3g_0^3] \quad (B 4)
\end{aligned}$$

$$\begin{aligned}
a_4 = & +63(1+e)\pi^{3/2}St_d\nu g_0 \left( 1543500000(59-23e)(1-e)\pi^2 \right. \\
& +52500\pi St_d^2[25(14647-16830e+4919e^2)\pi \\
& +140(1+e)(2041-6273e+5999e^2-1383e^3)\pi\nu g_0 \\
& -588(1-e)(1+e)^2\{360(3-e)^2(1+3e)+(1-3e)^2(59-19e)\pi\}\nu^2g_0^2] \\
& -20321280000(1-e)(3-e)^3(1+e)^4St_d^3\nu^3g_0^3-175\pi St_d^4[625(13-9e)(107+193e)\pi \\
& -500(1+e)(28181-94445e+71523e^2-14283e^3)\pi\nu g_0 \\
& +100(1+e)^2\{96(3-e)(1+3e)(4267-4873e+1134e^2) \\
& +(10223-137768e+356922e^2-219384e^3+36855e^4)\pi\}\nu^2g_0^2 \\
& +5760(3-e)(1+e)^3\{6(1736-439e-1557e^2+831e^3-123e^4) \\
& +7(1-3e)^2(1+3e)\pi\}\nu^3g_0^3-290304(27-7e)(1-3e)(1-e)(3-e)^2(1+e)^5\nu^4g_0^4] \\
& -38400(3-e)^2(1+e)^4St_d^5\nu^3g_0^3\{5(34879-39854e+9151e^2) \\
& +42(1+e)(519+41e-191e^2+15e^3)\nu g_0 \\
& -672(3-e)(1+e)^2St_d^6\nu^2g_0^2[625(1+3e)(382-391e+81e^2)\pi \\
& +375(1+e)(1303+394e-984e^2+198e^3-15e^4)\pi\nu g_0 \\
& -150(3-e)(1+e)^2\{16(1+3e)^2+(325-1476e-654e^2+1044e^3-135e^4)\pi\}\nu^2g_0^2 \\
& +1440(3-e)^2(1+e)^4(1+3e)\{6(1-e)-(1-3e)\pi\}\nu^3g_0^3 \\
& \left. -7776(1-e)^2(3-e)^3(1+e)^6\nu^4g_0^4 \right) \quad (B 5)
\end{aligned}$$

$$\begin{aligned}
a_5 = & +3780(3-e)(1+e)^2\pi St_d^2\nu^2g_0^2 \left( 463050000(11-5e)(1-e)\pi^2 \right. \\
& +4200\pi St_d^2[25(6405-8420e+3047e^2)\pi \\
& +140(1+e)(963-3040e+2881e^2-660e^3)\pi\nu g_0 \\
& -588(1-e)(1+e)^2\{120(3-e)^2(1+3e)+(29-9e)(1-3e)^2\pi\}\nu^2g_0^2] \\
& -304819200(1-e)(3-e)^3(1+e)^4St_d^3\nu^3g_0^3-7\pi St_d^4[625(13-9e)(131+205e)\pi \\
& -500(1+e)(12821-45203e+33483e^2-6381e^3)\pi\nu g_0 \\
& +100(1+e)^2\{48(3-e)(1+3e)(2873-3281e+756e^2) \\
& +(5147-66416e+173322e^2-104688e^3+16443e^4)\pi\}\nu^2g_0^2 \\
& +2880(3-e)(1+e)^3\{2(2556-235e-2253e^2+963e^3-135e^4) \\
& +7(1-3e)^2(1+3e)\pi\}\nu^3g_0^3-145152(1-3e)(13-3e)(1-e)(3-e)^2(1+e)^5\nu^4g_0^4] \\
& -768(3-e)^2(1+e)^4St_d^5\nu^3g_0^3\{5(8783-10054e+2279e^2) \\
& \left. +42(1+e)(127+21e-55e^2+3e^3)\nu g_0 \right\} \quad (B 6)
\end{aligned}$$

$$\begin{aligned}
a_6 = & +15876(3-e)^2(1+e)^3\sqrt{\pi}St_d^3\nu^3g_0^3 \left( 132300000(17-9e)(1-e)\pi^2 \right. \\
& +St_d^2[5000(32639-50458e+21923e^2)\pi^2 \\
& +252000(1+e)(605-1967e+1843e^2-417e^3)\pi^2\nu g_0 \\
& -352800(1-e)(1+e)^2\pi\{180(3-e)^2(1+3e)+(57-17e)(1-3e)^2\pi\}\nu^2g_0^2] \\
& -34836480(1-e)(3-e)^3(1+e)^4St_d^3\nu^3g_0^3 \\
& -\pi St_d^4[625(1951+1362e-1969e^2)\pi \\
& -500(1+e)(7757-28985e+20859e^2-3663e^3)\pi\nu g_0 \\
& +100(1+e)^2\{96(3-e)(1+3e)(724-829e+189e^2) \\
& +(3455-42632e+112122e^2-66456e^3+9639e^4)\pi\}\nu^2g_0^2] \\
& +1920(3-e)(1+e)^3\{3(1009+78e-864e^2+258e^3-33e^4) \\
& +7(1-3e)^2(1+3e)\pi\}\nu^3g_0^3 -145152(1-3e)(1-e)(3-e)^2(5-e)(1+e)^5\nu^4g_0^4] \quad (\text{B } 7)
\end{aligned}$$

$$\begin{aligned}
a_7 = & 3048192(3-e)^3(1+e)^4St_d^4\nu^4g_0^4 \left( 275625(47-29e)(1-e)\pi^2 \right. \\
& +5\pi St_d^2[25(3325-6284e+3211e^2)\pi+210(1+e)(427-1434e+1325e^2-294e^3)\pi\nu g_0 \\
& -1764(1-e)(1+e)^2\{18(3-e)^2(1+3e)+(1-3e)^2(7-2e)\pi\}\nu^2g_0^2] \\
& \left. -36288(1-e)(3-e)^3(1+e)^4St_d^3\nu^3g_0^3 \right) \quad (\text{B } 8)
\end{aligned}$$

$$\begin{aligned}
a_8 = & +2286144(1-e)(3-e)^4(1+e)^5\sqrt{\pi}St_d^5\nu^5g_0^5 \left( 264600(43-31e)\pi \right. \\
& +St_d^2[25(4069-5917e)\pi+420(1+e)(321-796e+219e^2)\pi\nu g_0 \\
& \left. -1764(1+e)^2\{24(3-e)^2(1+3e)+(1-3e)^2(11-3e)\pi\}\nu^2g_0^2] \right) \quad (\text{B } 9)
\end{aligned}$$

$$a_9 = +725896442880(13-11e)(1-e)(3-e)^5(1+e)^6\pi St_d^6\nu^6g_0^6 \quad (\text{B } 10)$$

$$a_{10} = +1451792885760(3-e)^6(1-e)^2(1+e)^7\sqrt{\pi}St_d^7\nu^7g_0^7. \quad (\text{B } 11)$$

Note that  $a_{10} = 0 = a_9 = a_8$  for  $e = 1$ , leading to a seventh-order polynomial for the temperature equation (3.1).

### B.1. Coefficients in equation (3.2)

The expressions for the coefficients of the quadratic polynomial (3.2) are given by

$$\begin{aligned}
\alpha = & \left\{ 4375(403507877743-2246328611260e+3804527790618e^2 \right. \\
& -2609043261628e^3+637175564527e^4)+7000(1841694494389-7655913090100e \\
& +9936315870558e^2-5002578505972e^3+804447129877e^4)\nu g_0 \\
& -1600(62781766140229-251533031265556e+383366453401950e^2 \\
& -256170205851796e^3+62747567827717e^4)\nu^2g_0^2-5120(15714866428013 \\
& -15498531439412e-53893773596274e^2+89513065956556e^3-35535888468883e^4)\nu^3g_0^3 \\
& +16384(55881178513091-252823502621588e+422083141225338e^2 \\
& \left. -306989989576772e^3+81830578488731e^4)\nu^4g_0^4 \right\}, \quad (\text{B } 12)
\end{aligned}$$

$$\begin{aligned} \mathfrak{b} = & -\left\{ 625(18778586124871 - 207069149476948e + 405023838452178e^2 \right. \\ & - 286513612682692e^3 + 68321253421567e^4) - 4000(195746263524613 \\ & - 832915077870466e + 1329037811913048e^2 - 934639544844094e^3 \\ & + 243370589968483e^4)\nu g_0 - 3200(1463725651004753 - 5379389667615500e \\ & + 7142510907700158e^2 - 4045727120864924e^3 + 816845472782489e^4)\nu^2 g_0^2 \\ & + 10240(1346413690308923 - 5798356073502980e + 9330745323357066e^2 \\ & - 6630820234286324e^3 + 1752390117615923e^4)\nu^3 g_0^3 \\ & + 262144(16659939237511 - 30369558221428e - 11077909546134e^2 \\ & \left. + 44890628459564e^3 - 20186596004777e^4)\nu^4 g_0^4 \right\}, \end{aligned} \quad (\text{B } 13)$$

$$\begin{aligned} \mathfrak{c} = & 48\left\{ 1875(5292277916157 - 7198799909266e - 8983084338480e^2 \right. \\ & + 18332139150594e^3 - 7387862080253e^4) - 1000(653424923400545 \\ & - 2603969210032424e + 3825404450727618e^2 - 2459579042010032e^3 \\ & + 584400338366645e^4)\nu g_0 + 2400(449794922964867 - 2000985299639524e \\ & + 3319509658053858e^2 - 2435905168315812e^3 + 667705483931683e^4)\nu^2 g_0^2 \\ & + 7680(1001122119985027 - 3916700322857816e + 5710226196969174e^2 \\ & - 3680145930163120e^3 + 885453528512687e^4)\nu^3 g_0^3 \\ & + 4096(1 - e)(1554567774124823 - 5288085618364101e \\ & \left. + 5986900650648069e^2 - 2247667963924055e^3)\nu^4 g_0^4 \right\}. \end{aligned} \quad (\text{B } 14)$$

## Appendix C. Ignited state: analysis in the principal axes frame

The second-moment balance equation (4.1) in the principal-axes frame can be written as four independent equations (Alam, Saha & Gupta 2019):

$$-4\eta\rho T\dot{\gamma}\cos 2\phi + 2\dot{\gamma}[(\Theta_{x'x'} - \Theta_{y'y'})\cos 2\phi - 2\Theta_{x'y'}\sin 2\phi] + \frac{6\rho T\dot{\gamma}}{St_d} = A_{\alpha'\alpha'}, \quad (\text{C } 1)$$

$$-4\eta\rho T\dot{\gamma}\cos 2\phi + 2\dot{\gamma}[(\Theta_{x'x'} - \Theta_{y'y'})\cos 2\phi - 2\Theta_{x'y'}\sin 2\phi] + \frac{12\rho T\dot{\gamma}\lambda^2}{St_d} = -3\hat{G}_{z'z'}, \quad (\text{C } 2)$$

$$4(1 + \lambda^2)\rho T\dot{\gamma}\cos 2\phi + 2\dot{\gamma}(\Theta_{x'x'} + \Theta_{y'y'})\cos 2\phi - \frac{4\rho T\dot{\gamma}\eta}{St_d} = \Gamma_{x'x'} - \Gamma_{y'y'}, \quad (\text{C } 3)$$

$$2\rho T\dot{\gamma}[\eta - (1 + \lambda^2)\sin 2\phi] - (\Theta_{x'x'} + \Theta_{y'y'})\dot{\gamma}\sin 2\phi = \Gamma_{x'y'}. \quad (\text{C } 4)$$

They represent (i) the trace of (4.1), with  $A_{\alpha'\alpha'} = A_{x'x'} + A_{y'y'} + A_{z'z'}$ , (ii) the  $z'-z'$  component of the deviatoric part of (4.1) (iii) the difference between the  $x'-x'$  and  $y'-y'$  components and (iv) the off-diagonal  $x'-y'$  component of (4.1). The expressions for various integrals  $A_{\alpha'\beta'}$ ,  $\hat{E}_{\alpha'\beta'}$ ,  $\hat{G}_{\alpha'\beta'}$  and  $\Theta_{\alpha'\beta'}$  appearing in (C 1-C 4) remain the same as those for the dry granular flow; these have been evaluated explicitly by Saha & Alam (2016).

Retaining terms up-to fourth-order  $O(\eta^i \lambda^j R^k \sin^l 2\phi)$ ,  $i+j+k+l \leq 4$  in the expressions of  $A_{\alpha'\beta'}$ ,  $\hat{E}_{\alpha'\beta'}$ ,  $\hat{G}_{\alpha'\beta'}$  and  $\Theta_{\alpha'\beta'}$ , the above equations (C 1-C 4) simplify to the following

set of four coupled nonlinear algebraic equations

$$\begin{aligned} & 5\sqrt{\pi}\eta R \cos 2\phi + 4(1+e)\nu g_0 R (\sqrt{\pi}\eta \cos 2\phi + 8R) \\ & - \frac{3}{4}(1-e^2)\nu g_0 (10 + \eta^2 + 32R^2 + 8\sqrt{\pi}\eta R \cos 2\phi) - \frac{15\sqrt{\pi}R}{St_d} \\ & - \frac{1}{28}\nu g_0 R^4 \left\{ 256 - 16(2 + \cos 4\phi)(\frac{\eta}{R})^2 + 192(\frac{\lambda}{R})^2 - 6(\frac{\eta}{R})^2(\frac{\lambda}{R})^2 + \frac{3}{4}(\frac{\eta}{R})^4 + 64(\frac{\lambda}{R})^4 \right\} \\ & + \frac{16}{21}(1+e)\nu g_0 R^4 \left\{ 32 - (2 + \cos 4\phi)(\frac{\eta}{R})^2 + 12(\frac{\lambda}{R})^2 \right\} = 0, \end{aligned} \quad (C5)$$

$$\begin{aligned} & 35\sqrt{\pi}\eta R \cos 2\phi + (1+e)\nu g_0 [32(1+3e)R^2 - 3(3-e)(\eta^2 + 21\lambda^2) \\ & - 8\sqrt{\pi}(4-3e)\eta R \cos 2\phi] - \frac{210\sqrt{\pi}\lambda^2 R}{St_d} \\ & + \frac{16}{33}(1+e)\nu g_0 R^4 \left\{ 32(5+3e) + 2(2 + \cos 4\phi)(2-3e)(\frac{\eta}{R})^2 \right. \\ & \left. - 33(5-3e)(\frac{\lambda}{R})^2 - \frac{9}{32}(3-e)[(\frac{\eta}{R})^4 - 11(\frac{\eta}{R})^2(\frac{\lambda}{R})^2 - 66(\frac{\lambda}{R})^4] \right\} = 0, \end{aligned} \quad (C6)$$

$$\begin{aligned} & 5\sqrt{\pi}R \cos 2\phi - (1+e)\nu g_0 [3(3-e)\eta + 2\sqrt{\pi}(1-3e)R \cos 2\phi] - \frac{10\sqrt{\pi}\eta R}{St_d} \\ & + 5\sqrt{\pi}\lambda^2 R \cos 2\phi - (1+e)\nu g_0 R^3 \left\{ \frac{8}{7}\sqrt{\pi}[(4-3e)(\frac{\lambda}{R})^2 - 8(1+e)] \cos 2\phi \right. \\ & \left. + \frac{1}{42}(\frac{\eta}{R})[64(4-3e) - 32(5+3e)\cos 4\phi - 3(3-e)(\frac{\eta}{R})^2 + 36(3-e)(\frac{\lambda}{R})^2] \right\} = 0, \end{aligned} \quad (C7)$$

$$\begin{aligned} & 5(\eta - \sin 2\phi) + 2(1+e)(1-3e)\nu g_0 \sin 2\phi - 5\lambda^2 \sin 2\phi \\ & - \frac{8}{7}(1+e)\nu g_0 R^2 \sin 2\phi \left\{ 8(1+e) - (4-3e)(\frac{\lambda}{R})^2 + \frac{4}{3\sqrt{\pi}}(5+3e)\frac{\eta}{R} \cos 2\phi \right\} = 0. \end{aligned} \quad (C8)$$

Note that equations (C5-C6) contain only “even” order terms in  $(\eta, \sin 2\phi, \lambda, R)$ , with the neglected terms being of order six and beyond; on the other hand, equations (C7-C8) contain only “odd” order terms in  $(\eta, \sin 2\phi, \lambda, R)$ , with the neglected terms being of order five and beyond. Therefore, equations (C5-C8) belong to the “super-super-Burnett” orders since they incorporate all terms up-to “quartic-order” in the shear-rate ( $R \sim \dot{\gamma}$ ).

### C.1. Exact solution at Burnett order for the whole range of density

Removing the terms within the second-brackets in (C5-C8), we obtain the Burnett-order equations as given in (4.5). The latter equations admit an exact solution, given by (4.10a) and (4.11), as discussed in §4.1.

### C.2. Perturbation solutions beyond Burnett order

The super-Burnett and super-super-Burnett order equations [(C5-C8)] are solved using a regular perturbation expansion around the exact Burnett-order solution (4.10a) and (4.11):

$$\left. \begin{array}{lcl} \eta & = & \eta^{(2)} + \varepsilon\eta^{(3)} + \varepsilon^2\eta^{(4)} \\ \lambda & = & \lambda^{(2)} + \varepsilon\lambda^{(3)} + \varepsilon^2\lambda^{(4)} \\ R & = & R^{(2)} + \varepsilon R^{(3)} + \varepsilon^2 R^{(4)} \\ \sin 2\phi & = & \sin^{(2)} 2\phi + \varepsilon \sin 2\phi^{(3)} + \varepsilon^2 \sin 2\phi^{(4)} \end{array} \right\}. \quad (C9)$$

In the above expressions  $\varepsilon \sim \dot{\gamma}$  and the superscript “2” corresponds to the “Burnett-order” solution [i.e. the closed form expressions (4.10a) and (4.11) as given in §4.1] and the superscripts “3” and “4” correspond to the corrections at third and fourth orders, respectively.

Plugging (C9) into corresponding third and fourth [(C5-C8)] order equations and

after performing some cumbersome algebra we obtain a null-solution at third-order

$$\eta^{(3)} = 0 = \lambda^{(3)} = R^{(3)} = \sin 2\phi^{(3)}. \quad (\text{C } 10)$$

The solutions at fourth-order are given by :

$$\begin{aligned} \eta^{(4)} = & - \left[ \left[ \sqrt{\pi}(1+e)\nu g_0 \cos 2\phi^{(2)} \{5 - 2(1+e)(1-3e)\nu g_0\} \{1024(5+3e)R^{(2)4} \right. \right. \\ & - 192(1+3e)R^{(2)2}(\eta^{(2)2} - 4\lambda^{(2)2}) - 9(1-e)(\eta^{(2)4} - 8\eta^{(2)2}\lambda^{(2)2} + 84\lambda^{(2)4}) \} \Big] \\ & - \left[ 8 \left\{ 5\sqrt{\pi}\eta^{(2)} \cos 2\phi^{(2)} + 2(1+e)\nu g_0 (8(1+3e)R^{(2)} - (1-3e)\sqrt{\pi}\eta^{(2)} \cos 2\phi^{(2)}) \right\} \right. \\ & \times \left\{ 210\sqrt{\pi}\lambda^{(2)2}R^{(2)} \cos 2\phi^{(2)} - 48(1+e)\sqrt{\pi}\nu g_0 R^{(2)} \cos 2\phi^{(2)} ((4-3e)\lambda^{(2)2} - 8(1+e)R^{(2)2}) \right. \\ & \left. \left. - 3(1+e)\nu g_0 \eta^{(2)} (32(1-3e)R^{(2)2} - (3-e)(\eta^{(2)2} - \lambda^{(2)2})) \right\} \right] \\ & \frac{168 \left[ \sqrt{\pi} \cos 2\phi^{(2)} \left\{ 2\sqrt{\pi} \cos 2\phi^{(2)} \{5 - 2(1+e)(1-3e)\nu g_0\} R^{(2)} - 3(1-e^2)\nu g_0 \eta^{(2)} \right\} \right.}{\left. \times \{5 - 2(1+e)(1-3e)\nu g_0\} + 2\nu g_0 \left\{ 3(1+e)(3-e) + 10\sqrt{\pi}R_{St}^{(2)} \right\} \left\{ 5\sqrt{\pi}\eta^{(2)} \cos 2\phi^{(2)} \right. \right.} \\ & \left. \left. + 2(1+e)\nu g_0 (8(1+3e)R^{(2)} - (1-3e)\sqrt{\pi}\eta^{(2)} \cos 2\phi^{(2)}) \right\} \right] \end{aligned} \quad (\text{C } 11)$$

$$\begin{aligned} \lambda^{(4)} = & \frac{1}{77616\lambda^{(2)}} \left( \left[ \frac{28(1+e)}{\{3(1+e)(3-e) + 10\sqrt{\pi}R_{St}^{(2)}\}} \left\{ 1024(5+3e)R^{(2)4} \right. \right. \right. \\ & + 96R^{(2)2} (2(2-3e)\eta^{(2)2} - 11(5-3e)\lambda^{(2)2}) - 9(3-e)(\eta^{(2)4} - 11\eta^{(2)2}\lambda^{(2)2} - 66\lambda^{(2)4}) \} \Big]_a \\ & - \left[ \frac{132}{\left\{ 3(1+e)(3-e) + 10\sqrt{\pi}R_{St}^{(2)} \right\} \sqrt{\pi} \{5 - 2(1+e)(1-3e)\nu g_0\} \nu^2 g_0^2 \cos 2\phi^{(2)}} \right. \\ & \times \left[ 35\sqrt{\pi}\eta^{(2)} \cos 2\phi^{(2)} + 8(1+e)\nu g_0 \left\{ 8(1+3e)R^{(2)} - (4-3e)\sqrt{\pi}\eta^{(2)} \cos 2\phi^{(2)} \right\} \right] \\ & \times \left[ 70\sqrt{\pi}\lambda^{(2)2}R^{(2)} \cos 2\phi^{(2)} + (1+e)\nu g_0 \left\{ 16\sqrt{\pi}R^{(2)} \cos 2\phi^{(2)} (8(1+e)R^{(2)2} - (4-3e)\lambda^{(2)2}) \right. \right. \\ & \left. \left. - 32(1-3e)\eta^{(2)2}R^{(2)} + (3-e)\eta^{(2)} (\eta^{(2)2} - 12\lambda^{(2)2}) \right\} \right]_b \\ & + \left[ \frac{1848\eta^{(4)} \left[ \frac{\sqrt{\pi} \{35 - 8(1+e)(4-3e)\nu g_0\} R^{(2)} \cos 2\phi^{(2)} - 6(3-e)(1+e)\nu g_0 \eta^{(2)}}{\{3(1+e)(3-e)\nu g_0 + 10\sqrt{\pi}\nu g_0 R_{St}^{(2)}\}} \right.}{\left. \left. + \frac{\{35\sqrt{\pi}\eta^{(2)} \cos 2\phi^{(2)} + 8(1+e)\nu g_0 (8(1+3e)R^{(2)} - (4-3e)\sqrt{\pi}\eta^{(2)} \cos 2\phi^{(2)})\}}{\sqrt{\pi} \{5 - 2(1+e)(1-3e)\nu g_0\} \cos 2\phi^{(2)}} \right] \right]_c \right) \end{aligned} \quad (\text{C } 12)$$

$$\begin{aligned}
R^{(4)} = & \frac{1}{42\sqrt{\pi}\{5 - 2(1+e)(1-3e)\nu g_0\} \cos 2\phi^{(2)}} \left[ -210\sqrt{\pi}\lambda^{(2)2} R^{(2)} \cos^{(2)} 2\phi \right. \\
& + 48(1+e)\sqrt{\pi}\nu g_0 R^{(2)} \cos^{(2)} 2\phi \left\{ (4-3e)\lambda^{(2)2} - 8(1+e)R^{(2)2} \right\} \\
& + 3(1+e)\nu g_0 \eta^{(2)} \left\{ 32(1-3e)R^{(2)2} - (3-e)\left(\eta^{(2)2} - 12\lambda^{(2)2}\right) \right\} \\
& \left. + 42\eta^{(4)} \left\{ 3(3-e)(1+e) + 10\sqrt{\pi}R_{St}^{(2)} \right\} \nu g_0 \right] \quad (C13)
\end{aligned}$$

$$\begin{aligned}
\sin 2\phi^{(4)} = & \frac{1}{21\sqrt{\pi}\{5 - 2(1+e)(1-3e)\nu g_0\}} \left[ 105\sqrt{\pi}\left(\eta^{(4)} - \lambda^{(2)2} \sin^{(2)} 2\phi\right) \right. \\
& \left. - 8(1+e)\nu g_0 \sin^{(2)} 2\phi \left\{ 4(5+3e)\eta^{(2)} R^{(2)} \cos^{(2)} 2\phi - 3\sqrt{\pi}\left((4-3e)\lambda^{(2)2} - 8(1+e)R^{(2)2}\right) \right\} \right] \quad (C14)
\end{aligned}$$

where  $R_{St}^{(2)} = R^{(2)}/\nu g_0 St_d$ . In absence of interstitial gas, i.e., in the limit of  $St_d \rightarrow \infty$ , the  $St$ -dependent terms in (C 11-C 14) disappear and we obtain the corresponding super-super-Burnett order solutions for the dry-granular flow – the resulting expressions match exactly with the solution provided by Saha & Alam (2016) for the uniform shear flow of dry granular fluid.

In the dilute limit, the solutions at fourth-order are given by:

$$\begin{aligned}
\eta^{(4)} = & \frac{1}{56\{3(5-e)(1+e)\eta^{(2)} + 10\sqrt{\pi}R_{St}^{(2)}(St \cos 2\phi + 2\eta^{(2)})\}} \left\{ (1+e)(27-11e)\eta^{(2)4} \right. \\
& \left. + 252(1-e^2)\lambda^{(2)4} + 560\sqrt{\pi} \cos 2\phi \eta^{(2)} \lambda^{(2)2} St R_{st}^{(2)} - 24(1+e)(13-5e)\eta^{(2)2} \lambda^{(2)2} \right\}, \quad (C15)
\end{aligned}$$

$$\begin{aligned}
\lambda^{(4)} = & \frac{7(1+e)}{17248\lambda^{(2)}\{3(3-e)(1+e) + 10\sqrt{\pi}R_{St}^{(2)}\}} \left\{ 132(19-11e)\lambda^{(2)4} \right. \\
& \left. - (13+3e)\eta^{(2)4} + 176\eta^{(2)2} \lambda^{(2)2} - 88(5+3e)\eta^{(2)} \eta^{(4)} \right\}, \quad (C16)
\end{aligned}$$

$$\begin{aligned}
R^{(4)} = & \frac{\nu}{560\sqrt{\pi}\eta^{(2)} \cos^{(2)} 2\phi} \left\{ 3(1-e^2)\left(\eta^{(2)4} - 8\eta^{(2)2} \lambda^{(2)2} + 84\lambda^{(2)4} + 56\eta^{(2)} \eta^{(4)}\right) \right. \\
& \left. - 560\sqrt{\pi} \cos 2\phi St R_{St}^{(2)} \eta^{(4)} \right\}, \quad (C17)
\end{aligned}$$

$$\sin 2\phi^{(4)} = \left\{ \eta^{(4)} - \lambda^{(2)2} \sin^{(2)} 2\phi \right\}. \quad (C18)$$

where  $R_{St}^{(2)} = R^{(2)}/\nu$ .

### C.3. Burnett-order equations for dilute gas-solid suspension: scaling arguments

Considering the dry granular limit ( $St \rightarrow \infty$ ) of a dilute gas-solid suspension, the Burnett-order second-moment balance equations are given by:

$$\left. \begin{aligned}
& 20\sqrt{\pi}\eta R \cos 2\phi - 3(1-e^2)\nu(10+\eta^2) = 0 \\
& 35\sqrt{\pi}\eta R \cos 2\phi - 3(1+e)(3-e)\nu(\eta^2 + 21\lambda^2) = 0 \\
& 5\sqrt{\pi}R \cos 2\phi - 3(1+e)(3-e)\nu\eta = 0 \\
& \eta - \sin 2\phi = 0
\end{aligned} \right\}. \quad (C19)$$

This admits an analytical solution of the form (Saha & Alam 2016)

$$\eta^2 = \frac{10(1-e)}{(11-3e)}, \quad \lambda^2 = \frac{20(1-e)}{7(11-3e)}, \quad \sin 2\phi = \eta, \quad R = \frac{3(3-e)(1+e)}{5\sqrt{\pi} \cos 2\phi} \frac{\sqrt{10}}{\sqrt{11-3e}} \nu \sqrt{1-e} \quad (\text{C } 20)$$

At leading order, we have

$$\eta = \frac{\sqrt{5(1-e)}}{2}, \quad \lambda = \sqrt{\frac{5(1-e)}{14}}, \quad \sin 2\phi = \frac{\sqrt{5(1-e)}}{4}, \quad R = \frac{6}{\sqrt{5\pi}} \nu \sqrt{1-e}, \quad (\text{C } 21)$$

as was predicted also by Richman (1989). It is easy to verify that the following quantities,

$$\eta, \lambda, \sin 2\phi, R/\nu, \sim \sqrt{1-e}, \quad (\text{C } 22)$$

are of the same order which holds for a sheared dilute granular gas. Equation (C 22) helps to prove a scaling between the Stokes number and the inelasticity.

*Proposition: The following scaling,*

$$St \sim \sqrt{1-e}, \quad (\text{C } 23)$$

*holds for a sheared dilute gas-solid suspension.*

To clarify (C 23), let us consider the second-moment balance equations for a dilute suspension at the Burnett-order:

$$20\sqrt{\pi}\eta R \cos 2\phi - 3(1-e^2)\nu(10 + \eta^2) - \frac{60\sqrt{\pi}R}{St} = 0 \quad (\text{C } 24a)$$

$$35\sqrt{\pi}\eta R \cos 2\phi + (1+e)\nu[-3(3-e)(\eta^2 + 21\lambda^2)] - \frac{210\sqrt{\pi}\lambda^2 R}{St} = 0, \quad (\text{C } 24b)$$

$$5\sqrt{\pi}R \cos 2\phi - 3(3-e)(1+e)\nu\eta - \frac{10\sqrt{\pi}\eta R}{St} = 0, \quad (\text{C } 24c)$$

$$5(\eta - \sin 2\phi) = 0. \quad (\text{C } 24d)$$

From (C 24c), we find

$$5\sqrt{\pi}R \cos 2\phi \sim R \sim \nu\sqrt{1-e} \quad \text{and} \quad 3(1+e)(3-e)\nu\eta \sim \nu\sqrt{1-e},$$

and hence

$$\frac{10\sqrt{\pi}\eta R}{St} \sim \frac{\sqrt{1-e}\nu\sqrt{1-e}}{\sqrt{1-e}} \sim \nu\sqrt{1-e},$$

which implies that all three terms in (C 24c) are of the same order. Similarly, from (C 24b) we have

$$35\sqrt{\pi}\eta R \cos 2\phi \sim \nu(1-e) \quad \text{and} \quad 3(1+e)(3-e)\nu(\eta^2 + 21\lambda^2) \sim \nu(1-e),$$

and hence

$$\frac{210\sqrt{\pi}\lambda^2 R}{St} \sim \frac{(1-e)\nu\sqrt{1-e}}{\sqrt{1-e}} \sim \nu(1-e),$$

again implying that three terms in (C 24b) are of the same order. From the above ordering arguments the proposition made in (C 23) is justified.

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